

Magnetospheric Visualization

How to Extract Global Measurements for End Users

Asher Pembroke

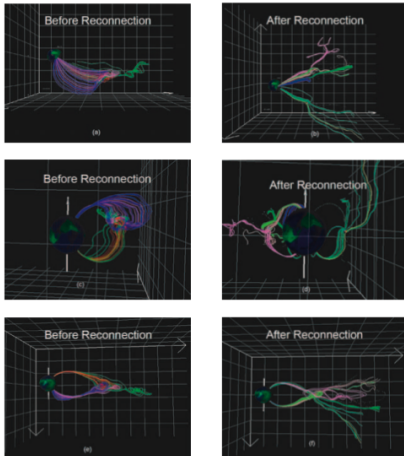
Rice University, Houston

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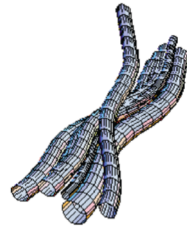
- 1 Introduction/Motivation
 - Previous Work in our Field
 - Unsolved Mysteries of the Magnetosphere
- 2 A Possible Solution
 - A New* Data Structure
 - Simple Examples
- 3 Workflow of Global Measurement
- 4 Implementation/What's to come

*See graphics literature for a similar approach to analysis for fluid flows [1].

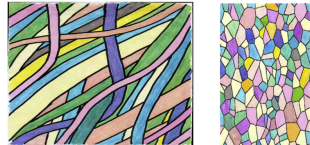
Examples from the Space Physics Community



D. Cai *et al.*, 2007



Consolini *et al.*, 2002



Problems with standard tools

- 2D Cross sections of scalars, including isocontours
- 3D field line traces through vector fields
- 3D isosurfaces (e.g. pressure, density)
- Artists' interpretations of the underlying structure

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- 2D Cross sections of scalars, including isocontours
 - 3D field line traces through vector fields
 - 3D isosurfaces (e.g. pressure, density)
 - Artists' interpretations of the underlying structure
- Existing Visualization techniques tend to *hide* the information scientists are interested in!
- Removes 3D structure. No natural point of reference.
 - Spaghetti plots. "But field lines don't exist!" arguments
 - Useful, but difficult to interpret. Fundamentally local.
 - Far removed from simulation...

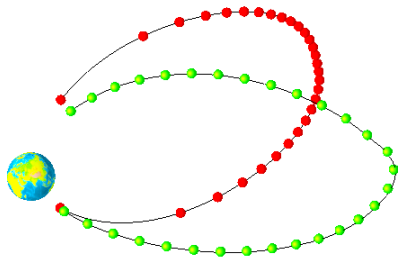
Open Questions

How do we reduce the simulations in a way that provides clear physical insight to users?

To Magnetospheric physicists, this means...

- How do we illustrate magnetic field geometry apart from field line depiction?
- How to tie field line geometry to global metrics?
- How do we avoid the overhead of redundancy in global field line traces?
- How do we interpret the mapping of ionospheric quantities, such as Field-Aligned Currents (FACs)?

Start Simple: Field line interpolation (1-D)



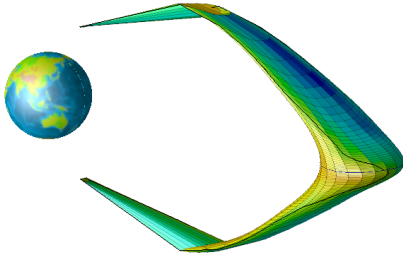
Field lines with resolution lowered to just 30 vertices each. Basic shape can still be captured.

Reduce the resolution of each field line.

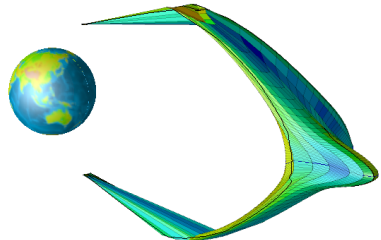
- Given $\vec{F}(\vec{r})$, can parameterize by fixed partially integrated $\int ds/F$.
- Drastically reduces storage requirements for each trace.
- Weighting schemes are flexible: Figure shows partial flux tube volume (red) and field line length to parameterize two different field lines.

Permits... Flux Tube Construction (2-D)

Using the previous result, we can build surfaces by tracing from a given curve's vertices.

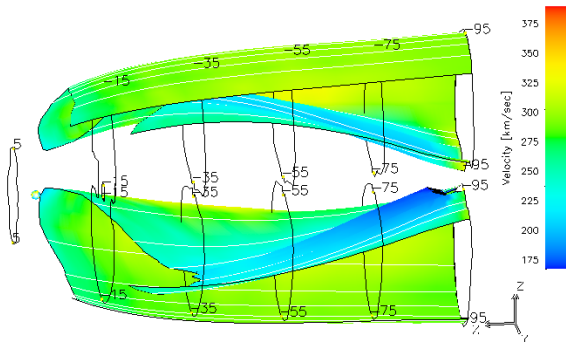


Flux tube mapped from a $1R_e$ ring of points in the equator (LFM)



The same flux tube, from a streakline trace after 8 minutes.

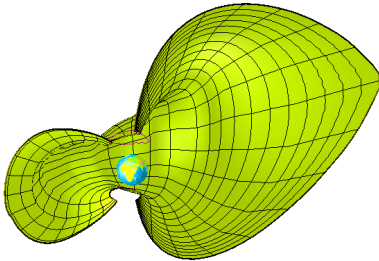
... Flow Surface Construction (2-D)



Plasma mantle drift surface, mapped backward in time from a $\beta = 1$ curve, $-95R_E$ in the tail.

The Field Line Volume (3-D): Mapping Electric Potential

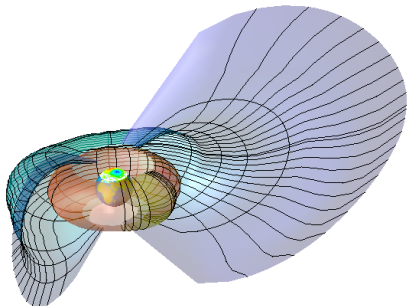
Volume elements can be built directly from surface connections.



LFM equipotential surface, created from field lines mapped from ionospheric surface elements.

- Yields a volumetric map of any quantity known at the surface.
- Previous Surfaces are now implicit - can be computed from isosurfaces.
- Implicit surface parameterizations (black curves).

The Field Line Volume (3-D): Drift Shell Construction



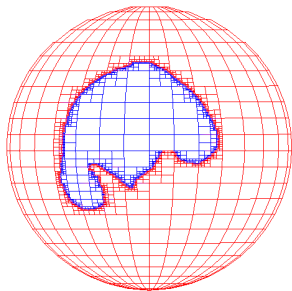
electron drift shells conserving
energy invariant (assuming $\frac{\partial B}{\partial t} = 0$)

$$v_D = \frac{\vec{B}}{B^2} \times \nabla \left(\frac{\lambda V^{-2/3}}{q} + \Phi \right)$$

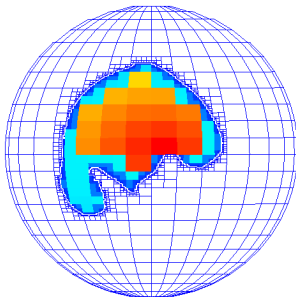
For energy invariant λ and flux tube
volume $V = \int ds/B$
Average particle drift: along
isosurfaces of an *effective* potential.
Drift shell visualization requires Φ
and $V^{-2/3}$ everywhere.

3D Global Separatrix: Start with Local Topology Search

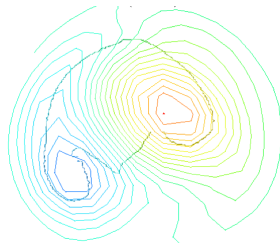
Multi-resolution Separatrix search
inspired by Warren & Schaefer's
scheme for subdivision surfaces [2]



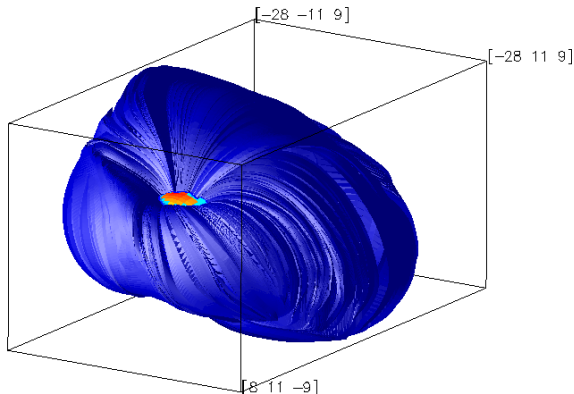
Fluxes computed per cell on the
starting mesh



Reconnection potential can be
calculated from $V_{max} - V_{min}$ along
the separatrix



3D Global Separatrix: ... Build the Field Line Volume



3D separatrix surface, interpolated from 2602 field lines forming a volume. Compare with brute force volume search (61k).

Bounding box gives us a measure of both the geoeffective length and magnetopause location!

A framework for developing new metrics

A Suggested Workflow...

- (1) Build a frame out of existing field line data (the bottleneck)
- (2) Populate it with global physical measurements - your field integrations, topologies, ionospheric quantities, etc.
- (3) Reparameterize to allow user definitions for regions of interest (ROI)
- (4) Extract your characteristic measurement
- (5) Apply to time series analysis if possible

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- (6) Make bold statements about the physics captured by your code! (optional)

Defining your ROI

Choose wisely...

- use what you already know about your region of interest to guide your intuition! Ex: if entropy is conserved along along drift paths, what does that tell us about the motion of a constant entropy surface? How is the frozen-in condition *measured*?
- Build the ROI so that integrations are convenient: Boundary normals should point out of the ROI, and line integrals should be defined on the boundary surface and follow a right-hand rule (can choose field and surface normal).

Making your measurement

Could be the ROI's enclosed mass, volume, energy, or momentum.

Examples:

- the structure of FACs
- the volume of a flux tube volume
- the energy within a flux tube volume
- the total forces acting on a ROI and the relative forces
- the magnetic flux into the ROI ($\int E \cdot d\vec{l}$)
- the integrated vorticity of the flow ($\int V \cdot d\vec{l}$)

Tracking Measurement with Time

Suggestions & Words of warning

- Use streaklines or use quantities that we know are conserved with the flow.
- Be careful when applying the maxwell-faraday equation - it only works when the integration surface is tied to the flow!
- Exploit the divergence-less nature of your fields - can use half the integrated flux magnitude for closed surfaces, and use the zero-flux line for path integrations.

Implementation

- A set of macros/networks built in CISM-DX (Open-DX) with rapid prototyping.
- Project hosted on google code:
<http://code.google.com/p/cismdxalgs/>
- Repository has video tutorials and a rough draft of a paper in progress. Comments/suggestions? apembroke@gmail.com
- AR Toolkit...

Delivery to the Community

There is a lot to do over the coming months...

- Integration with the CCMC framework will occur over the next year (starting now).
- Need to port methods into more stable working environments.
- Development of sanity checks are crucial: do ROI fluxes match volume integrals? Does the geometric volume match the volume within a flux tube?
- How do we adapt the approach to other Heliospheric regions?
- How do we convert global measurements into something relevant for spacecraft?
- The calculus of differential forms (DEC methods) inspired the approach - we should probably take geometry more seriously!

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You will see greater things than these!



Acknowledgements

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Acknowledgements

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-  J. P. M. Hultquist, *Constructing stream surfaces in steady 3d vector fields*, VIS '92: Proceedings of the 3rd conference on Visualization '92 (Los Alamitos, CA, USA), IEEE Computer Society Press, 1992, pp. 171–178.
-  Joe Warren and Scott Schaefer, *A factored approach to subdivision surfaces*, IEEE Comput. Graph. Appl. **24** (2004), no. 3, 74–81.